

October 24, 2003

Dept of Electrical and Computer Engineering
at Rutgers University
the **Wireless Information Networks LAB**oratory

WINLAB
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Electromagnetic Channels
Write or Radiate: Inscribed Mass vs.

- **Interference Avoidance, Pricing & Spectrum Management**
 - Interference hurts \Leftrightarrow deal with it!
- **Channel Quality**
 - How good can that RF channel be? \Leftrightarrow really good!
- **Infostations:**
 - Delay tolerant? \Leftrightarrow transmit when near base!

10 Years of WINLAB Research (Infostations redux)

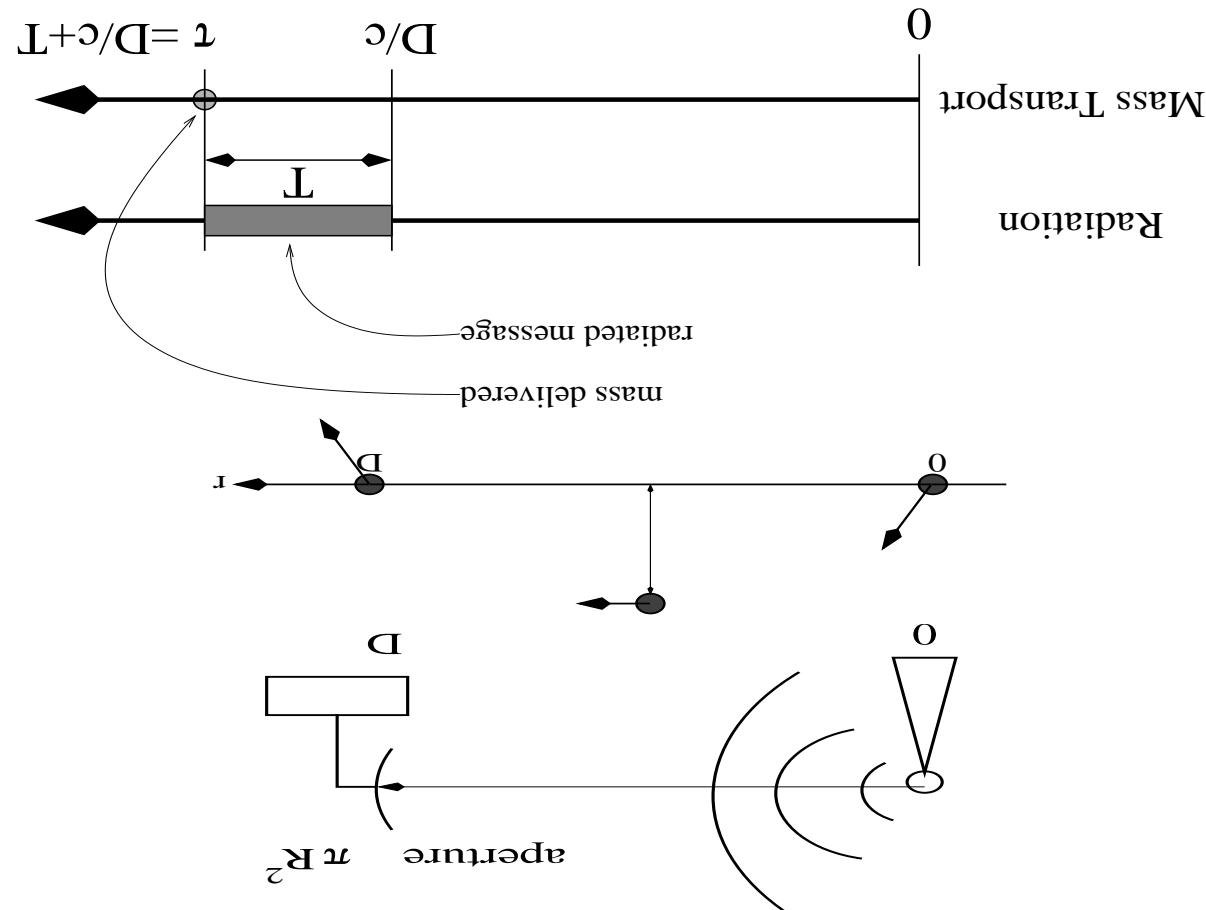
Completely ridiculous right?

- Forget RF! Write message down! Toss it to recipient!
- IMPLICATIONS:
- Can tolerate delay
- Channel good when nearby
- Storage density is increasing
- RF Interference is bad

A Hairy Old Epiphany!

- And maybe a LOT more room at the bottom
 - 1 bit per $\text{nm}^3 \rightarrow 1 \text{mm}^3 = 10^{18} \text{ bits!}$
- **RNA**: $1.8 \times 10^{24} \text{ bits/kg}$
- **STM** with Xe on Ni: $1.74 \times 10^{22} \text{ bits/kg}$
- E-beam **Lithography** with SiO_2 : $1.54 \times 10^{21} \text{ bits/kg}$
- Optical **Lithography** with SiO_2 : $3.85 \times 10^{18} \text{ bits/kg}$

A Little Empirical Rigor



A Little Analytic Rigor

- $\delta = \frac{D/\tau}{c}$: ratio of τ to the light travel time.
- T : radio messaging time (s).
- t : message deadline (s)
- N_0 : background noise energy (W Hz^{-1}).
- D : distance to receiver (m).
- $A = \pi R^2$: effective receiver aperture (m^2).
- W : bandwidth available for radiated communication (Hz).
- $\tilde{\rho}$: mass information density for inscribed information (bits kg^{-1}).
- B : message size (bits).

Parameters and Definitions

$$(\underline{A}) h \bar{\leq} [(\Lambda) h] E$$

- If $h()$ convex (Jensen):

$$[(\Lambda) h] E = (\Lambda) h^*(v)$$

- If V deterministic:

$$[(\Lambda) h] E \bar{\leq} (\Lambda) h^*(v)$$

- Max bigger than mean:

Rocket Science Foundations

with equality iff $v(t)$ is constant

$$(\underline{A}) h \leq [((v(t)) \max_{t \in [0, T]} h(v(t)) \geq \min_{t \in [0, T]} E[h(v(t))]$$

- Jensen says

$$\text{subject to } v = \frac{\dot{x}}{D}.$$

$$E_* = \min_{t \in [0, T]} \max_{v(t)} h(v(t))$$

- Minimum imparted energy

$$[(v(t)) E = \frac{\dot{x}}{D} = \int_t^0 v(t) dt = \frac{1}{2} \int_t^0 v^2(t) dt]$$

- Average velocity

Rocket Science

$$E_* \approx \frac{1}{2} m v^2$$

• $h(v) \approx \frac{1}{2} m v^2$:

$$\left(1 - \frac{\sqrt{1 - \frac{v^2}{c^2}}}{1} \right) E_* = mc^2$$

• $h(v) = mc^2$

$$\left(1 - \frac{\sqrt{1 - \frac{v^2}{c^2}}}{1} \right) E_* = mc^2$$

$$(v)h = E_*$$

• GIVEN: $h()$ and v

Minimum Transport Energy

$$0 = (x) \cdot b - (x) \cdot h''$$

$$0 = \frac{x\varrho}{\mathcal{E}\varrho} - \left(\frac{\alpha\varrho}{\mathcal{E}\varrho} \right) \frac{dt}{p}$$

- Calculus of variations:

$$E_* = \min_{\mathcal{E}} \max_{x(t)} \mathcal{E}(t) \leq \min_{\mathcal{E}} \int_0^T \frac{1}{2} \dot{x}(t)^2 dt$$

- Energy minimization:

$$((t)x)b + ((t)\alpha)h = (\mathcal{E}(t)$$

- $y(x)$ potential energy:

Potential Fields

- $E(t)$ constant \rightarrow minimization satisfied with equality, so ...
- Freefall? $\rightarrow E(t) = \text{constant}$
- $y(x)$ is force at position x : \rightarrow "free fall"
$$m\ddot{x} = y(x)$$
- Low speed:

Potential Field Results



- $E(t)$ constant \rightarrow minimization satisfied with equality, so ...
- Freefall? $\rightarrow E(t) = \text{constant}$
- $y(x)$ is force at position x : \rightarrow "free fall"
$$m\ddot{x} = y(x)$$
- Low speed:

Potential Field Results

- Pay a factor of 2 over free space

$$g_* = c\sqrt{2/gD}$$

- Delay at minimum energy
- Let $\delta = ct/D$.

$$E_* = \frac{1}{2}mgD$$

- Minimum energy:

Artillery Problem

- Escape examples (rough):
 - Earth: $\dot{q} < 2 \times 10^4$
 - Solar: $\dot{q} < 2 \times 10^3$
 - Milky Way: $\dot{q} < 6 \times 10^2$
- About a factor of 2 energy penalty, again
- Boils down to: need initial velocity larger than escape.
- Needs numerical calculation

Escape Problem

- Can (usually) ignore relativity
- Low speed ain't so low!
- Off by only $\approx 10\%$ at $0.4c$ and $\approx 50\%$ at $0.75c$

$$E_w \approx \frac{1}{B} \frac{\tilde{p}}{c^2} \left(\frac{c}{\tilde{q}} \right)^2$$

- $\tilde{q} \ll 1$:

$$E_w = \frac{\tilde{p}}{B c^2} \left(1 - \frac{\sqrt{\tilde{q}^2 - 1}}{\tilde{q}} \right)$$

- General

- Message size B , mass information density \tilde{p}

Inscribed Mass Energy Requirements

$$E_r \geq BN_0 \frac{A}{4\pi D^2} \ln 2$$

- Large TW :

$$E_r = BN_0 \frac{A}{4\pi D^2 TW} \left[2^{\frac{B}{TW}} - 1 \right]$$

- $E_r = PT$,

$$B = TC = TW \log_2 \left(\frac{4\pi D^2 N_0 W}{PA} + 1 \right)$$

- Bits a la Shannon:

$$V(D) = \frac{4\pi D^2}{A}$$

- Energy capture (no gain yet):

Radiation Energy Requirements

$$\sigma \leq \left[\frac{4\pi D^2}{\rho N_0} \right] \left[\frac{A}{c^2} \right] (2 \ln 2) \delta^2$$

- Large bandwidth W , $\delta \ll 1$

$$\sigma = \frac{E_w}{E_r}$$

- Definition:

Radiation to Transport Energy Ratio

Aperture	p_*	Far Field
(0.05 m)	$2.15 \times 10^{36} \left[\frac{\text{D}\delta}{\text{meter}} \right]^{-2} \frac{\text{bits}}{\text{kg}}$	$D = 0.37 \text{ m}$
(Arecibo)	$1.95 \times 10^{20} \left[\frac{\text{D}\delta}{\text{light year}} \right]^{-2} \frac{\text{bits}}{\text{kg}}$	$D \approx 3 \times 10^{-10} \text{ LY}$
(Earth)	$6.38 \times 10^{38} \left[\frac{\text{D}\delta}{\text{light year}} \right]^{-2} \frac{\text{bits}}{\text{kg}}$	$D \approx 0.6 \text{ LY}$

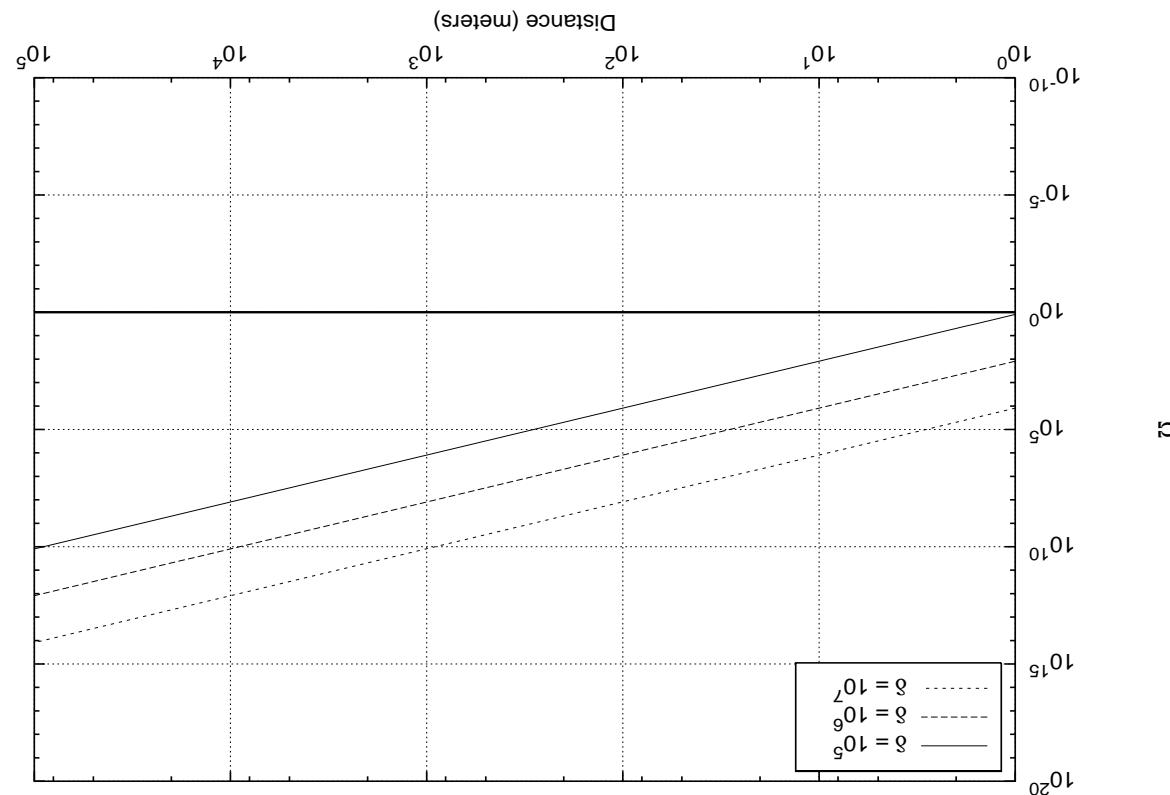
- p_* for $\alpha = 0.03m$:

- Critical p : $p_* = \arg_p \left\{ \frac{E_w(p)}{E_r(G)} \right\} = 1$

- Gain: $G_{\max} = \frac{8\pi^2 R^2}{\lambda^2}$

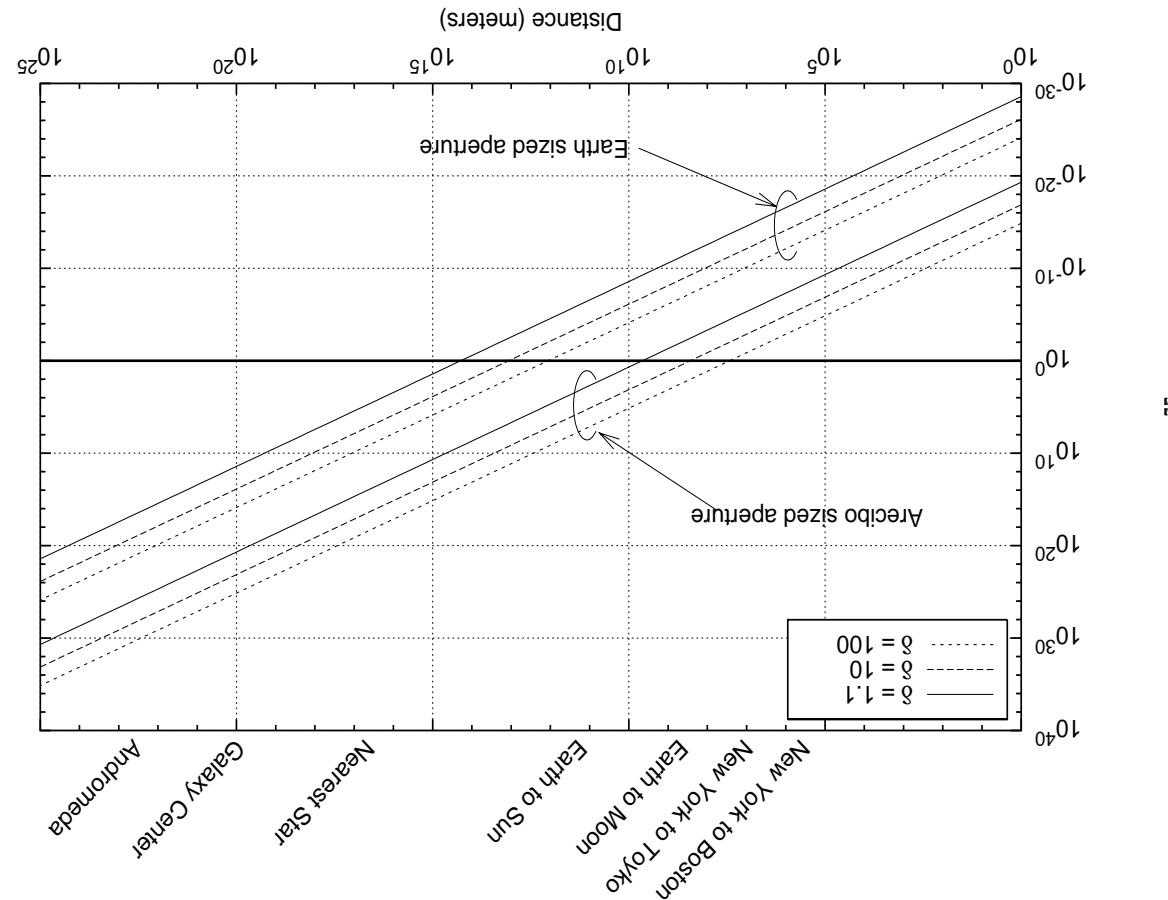
What About Gain?

0.05m receive aperture radius, 300°K, $\rho = 1.8 \times 10^{24} \text{ bits kg}^{-1}$



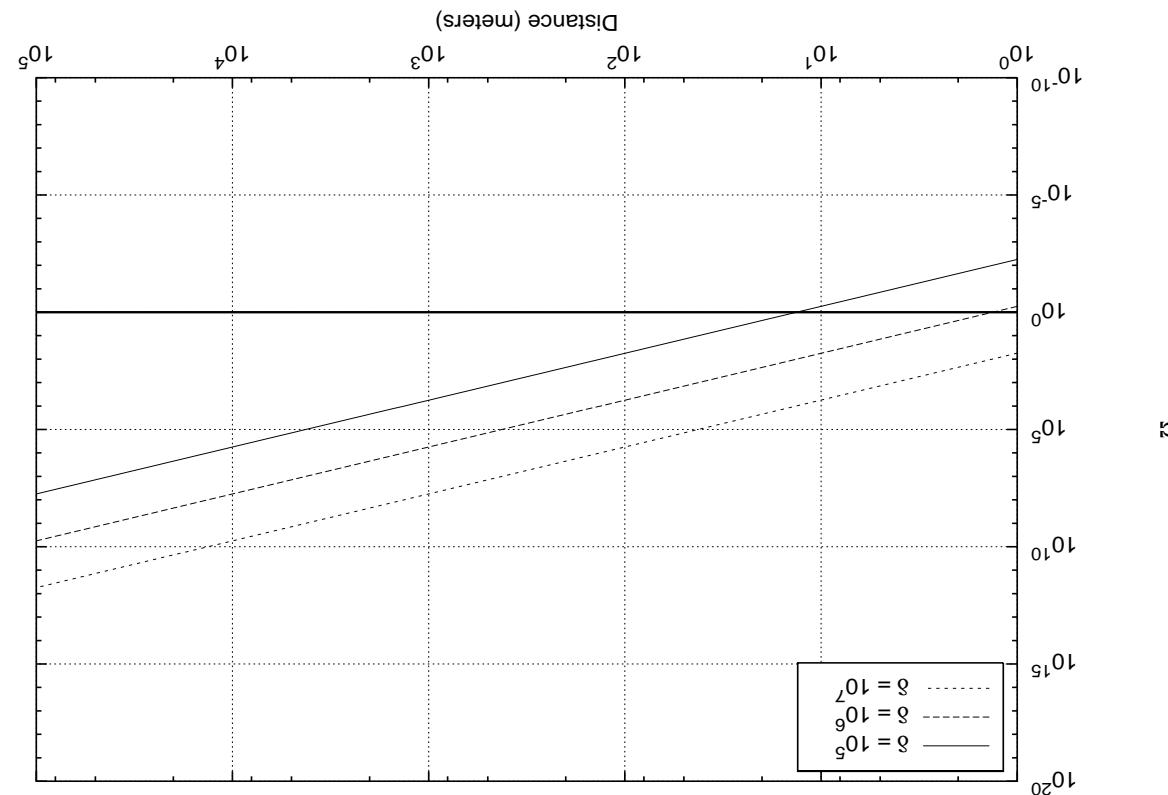
Q^2 vs. Distance: point to point no gain

$$3 \text{ K, } p = 1.8 \times 10^{24} \text{ bits kg}^{-1}$$



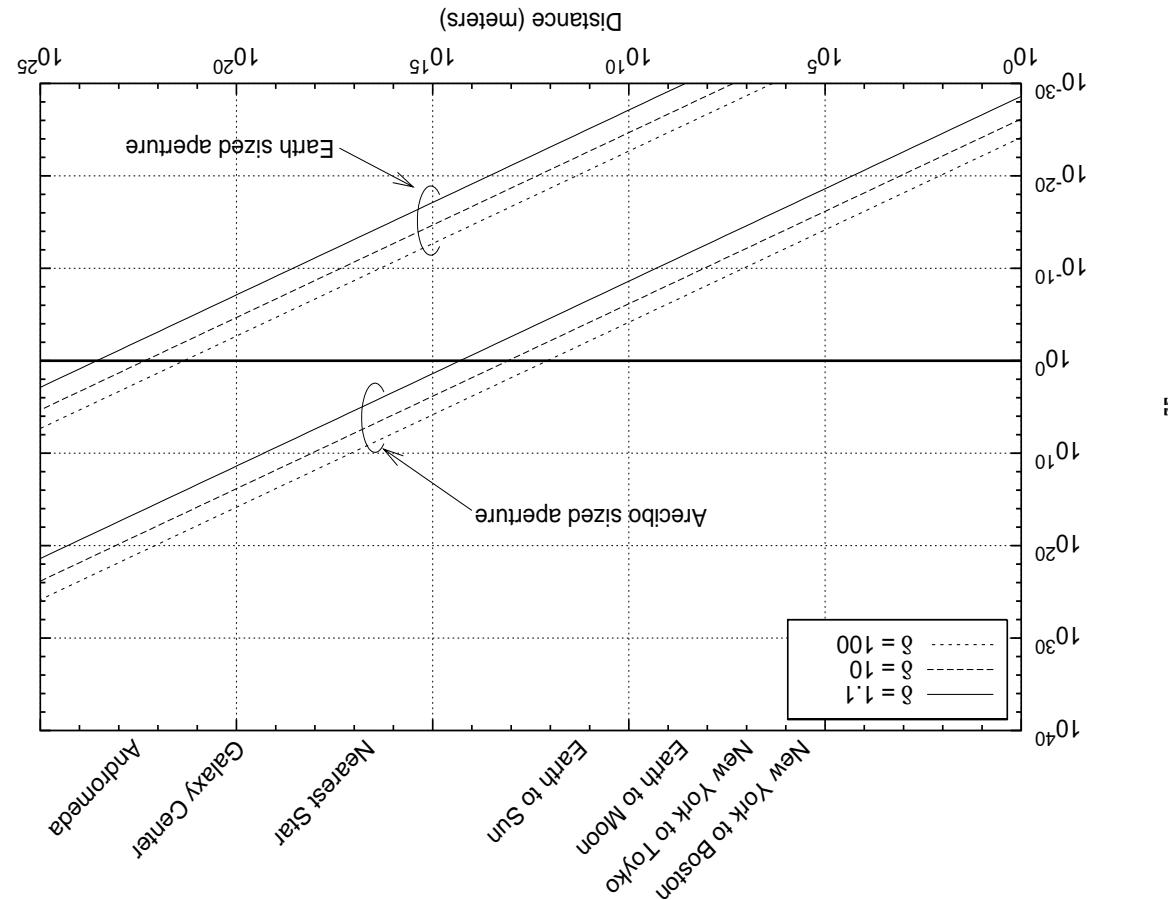
G vs. Distance: point to point no gain

0.05m receive aperture radius, 300°K, $\rho = 1.8 \times 10^{24} \text{ bits kg}^{-1}$



Q vs. Distance: point to point with gain

$$3 \text{ K, } g = 1.8 \times 10^{24} \text{ bits kg}^{-1}$$



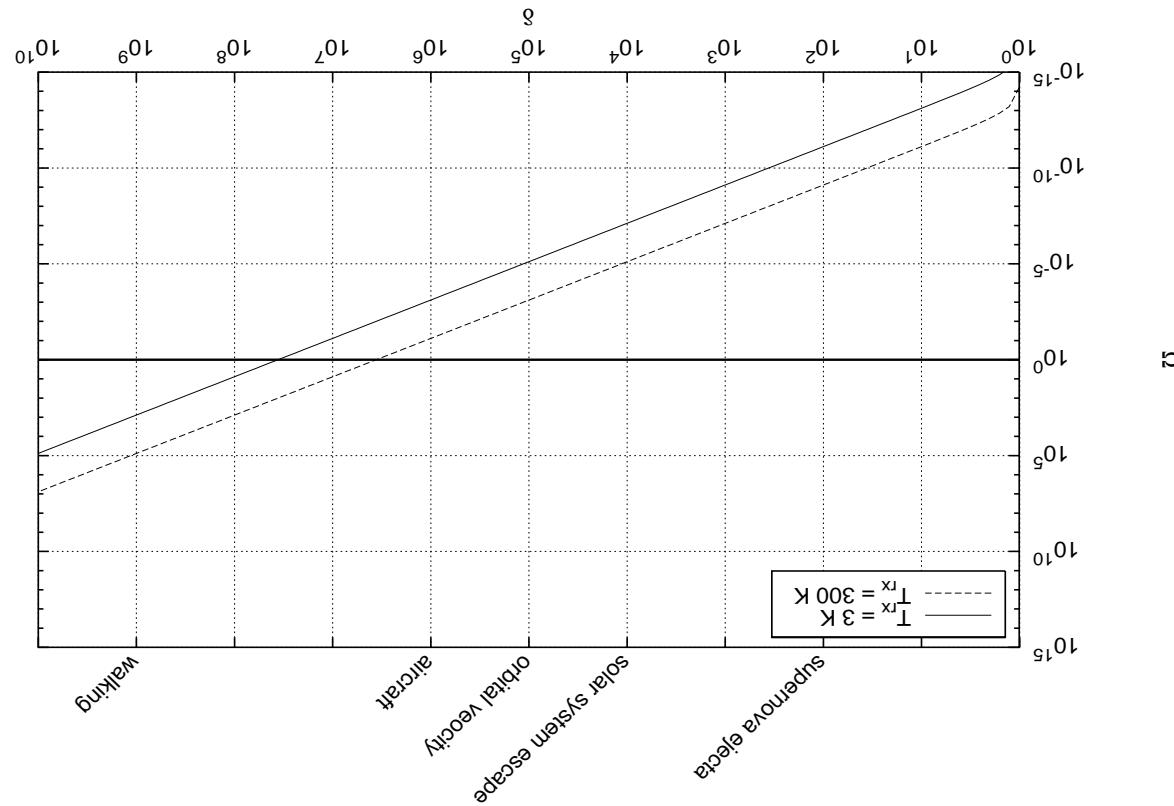
Q vs. Distance: point to point with gain

- Large BW limit: separate radiative channels (essentially pt2pt)
 - different messages, different receivers
- Multicast
 - one message, many receivers, known positions
 - Directed Broadcast
 - one message, many receivers, unknown positions
 - Blind Broadcast

What About Broadcast?

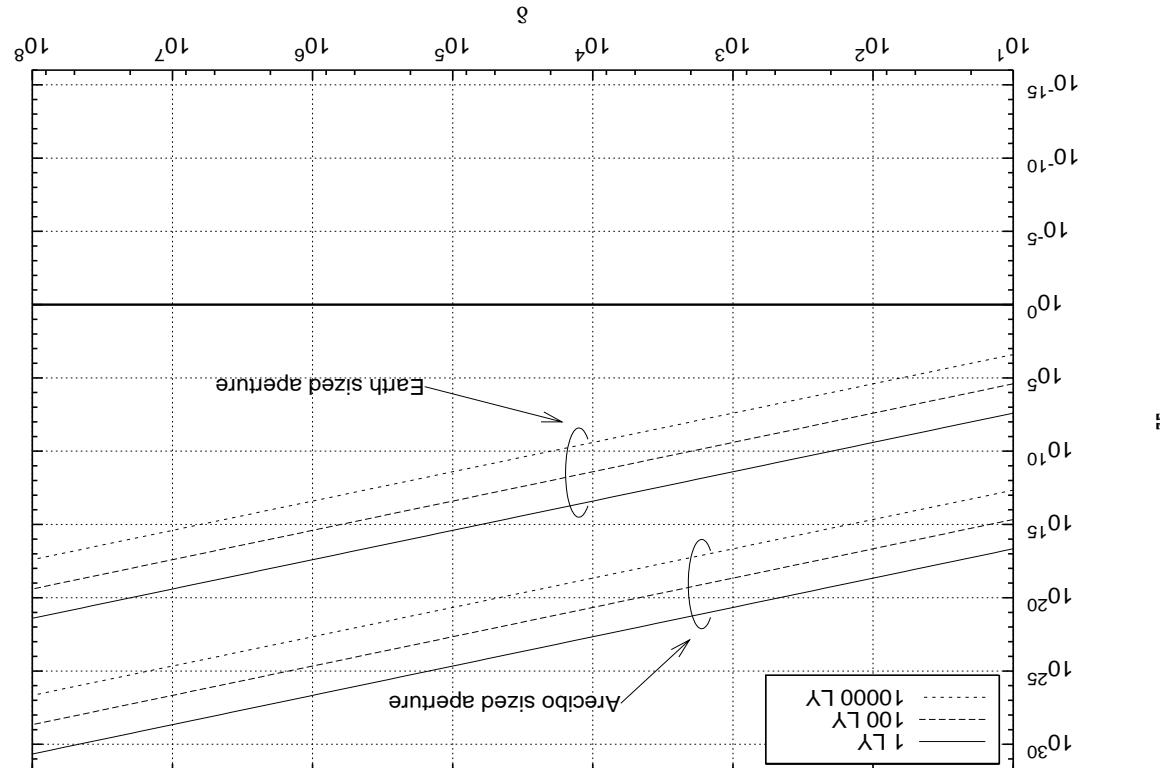
kg^{-1} and mass sucks unless you're a snail

Energy ratio vs. g . Receiver temperatures 30K and 300K, $\rho = 1.8 \times 10^{24}$ bits



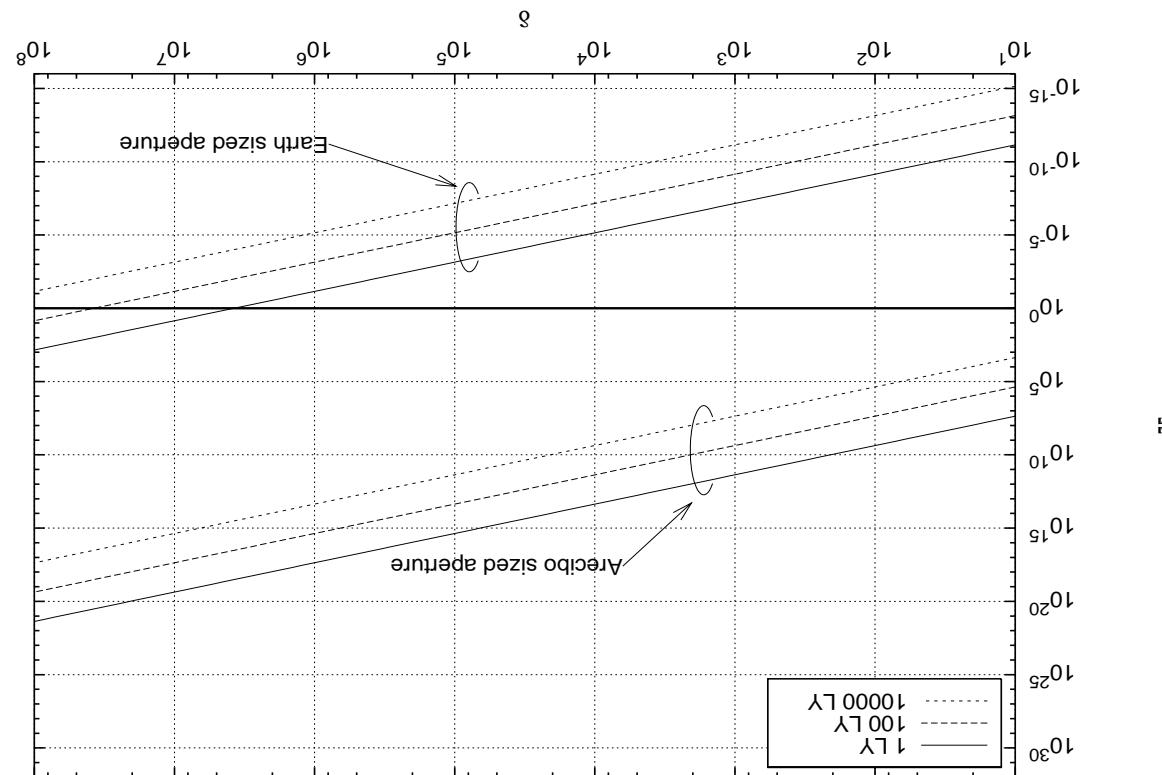
One Message, Unknown Locations

$\Omega = 6.4 \times 10^{-3} \text{ light year}^{-3}$ (stellar density of Milky Way).
 Energy ratio vs. Ω . $\rho = 1.8 \times 10^{24} \text{ bits kg}^{-1}, 3^\circ \text{K}$, receiver density



One Message, Known Locations, No Gain: interstellar

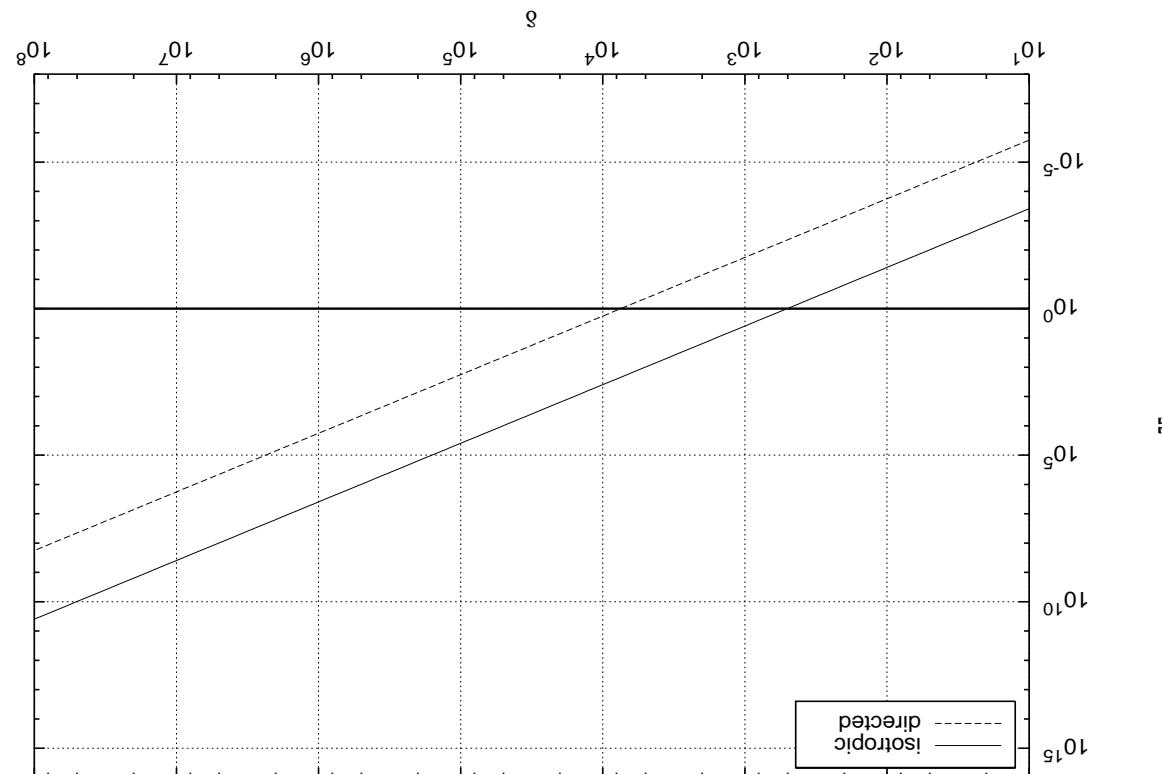
Same but with gain.



One Message, Known Locations with Gain: interstellar

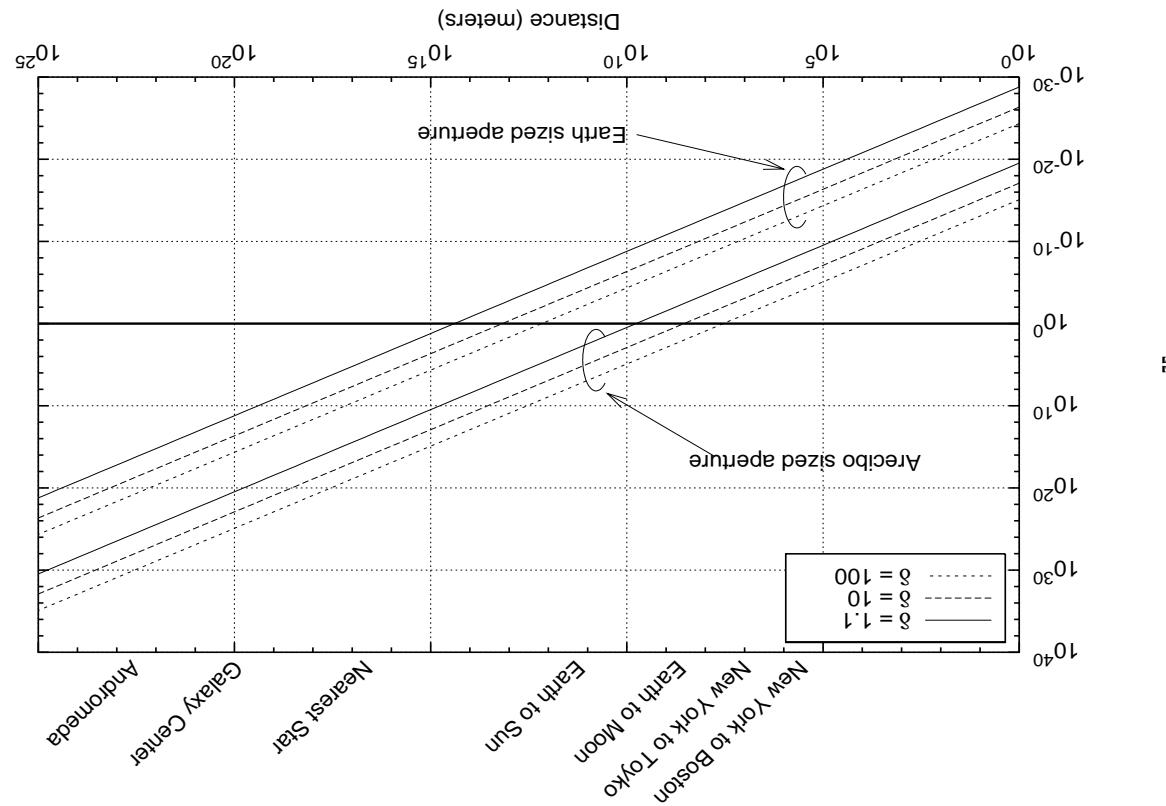
0.05 m, receiver density $\sigma = 0.01 \text{ m}^{-2}$.

Energy ratio vs. δ . $\rho = 1.8 \times 10^{24} \text{ bits kg}^{-1}$, 300°K, receiver aperture radius



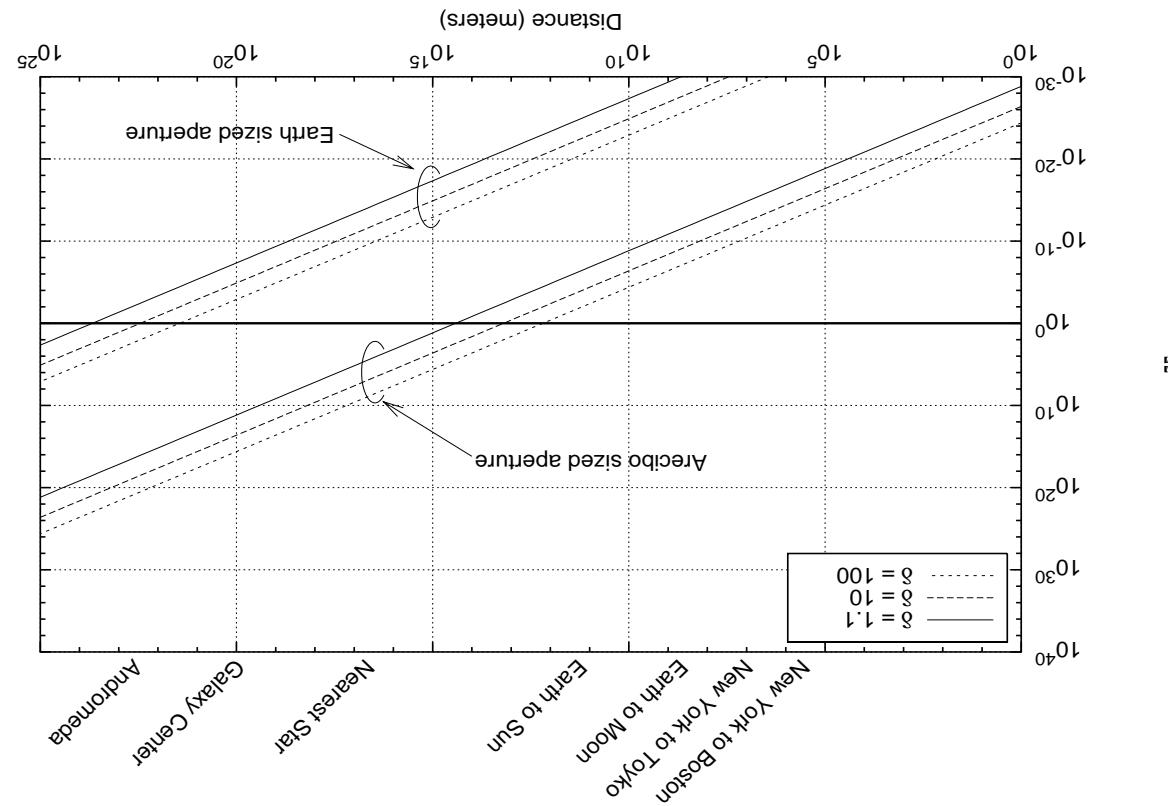
One Message, Known Locations: terrestrial

$\rho = 1.8 \times 10^{24} \text{ bits kg}^{-1}$, receiver temperature 3 K .



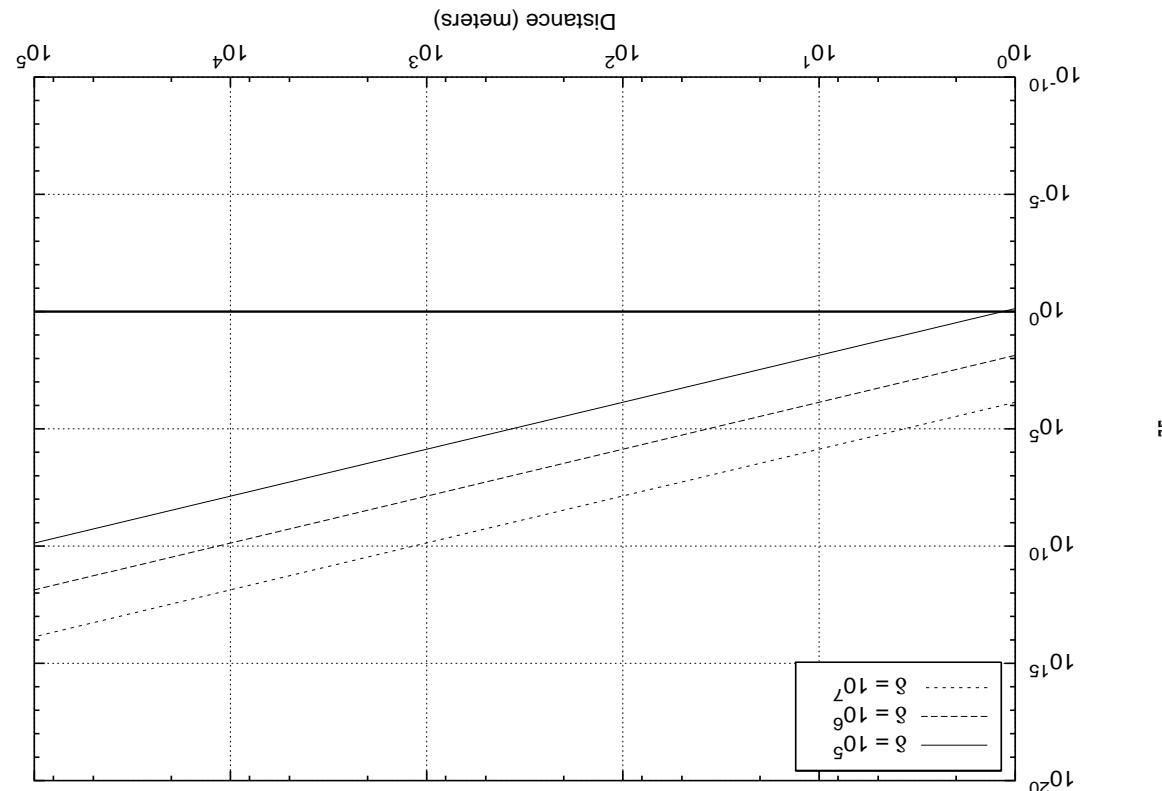
Many Messages, No Gain: interstellar

Same but with gain.



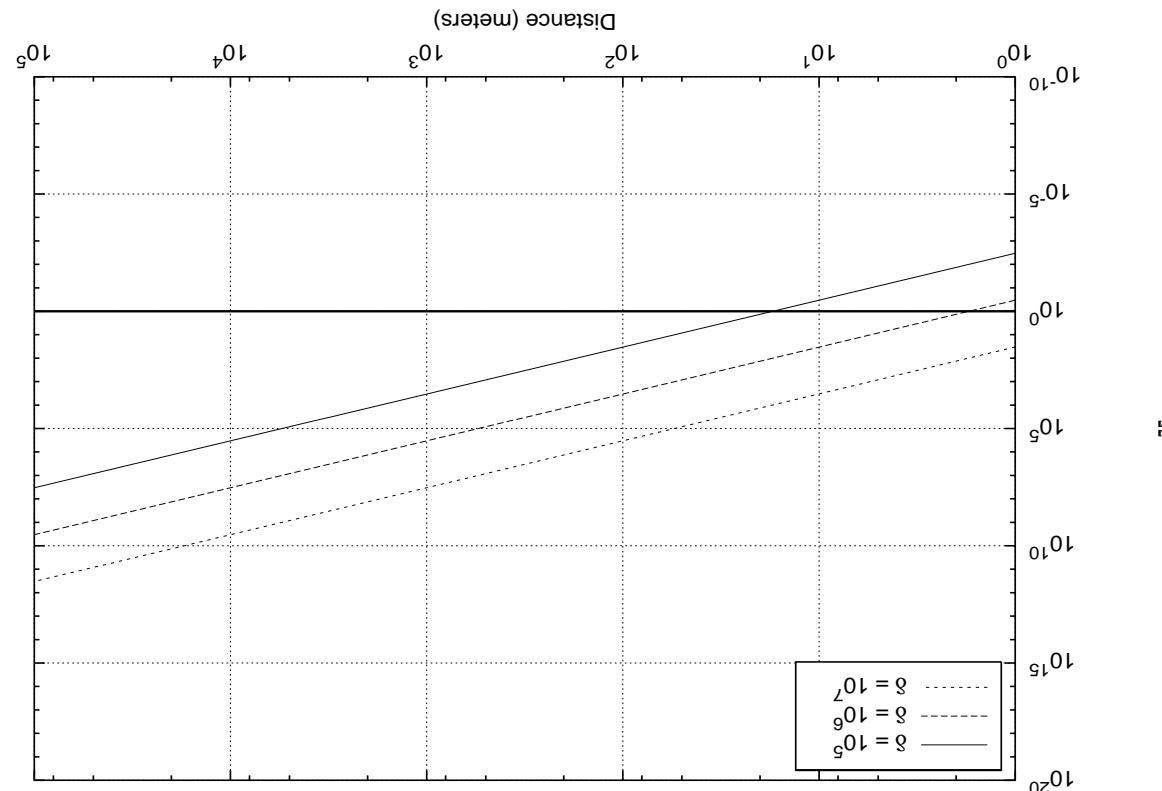
Many Messages with Gain: interstellar

Aperture 0.05 m, $\rho = 1.8 \times 10^{24} \text{ bits kg}^{-1}$, 300°K.



Many Messages, No Gain: terrestrial

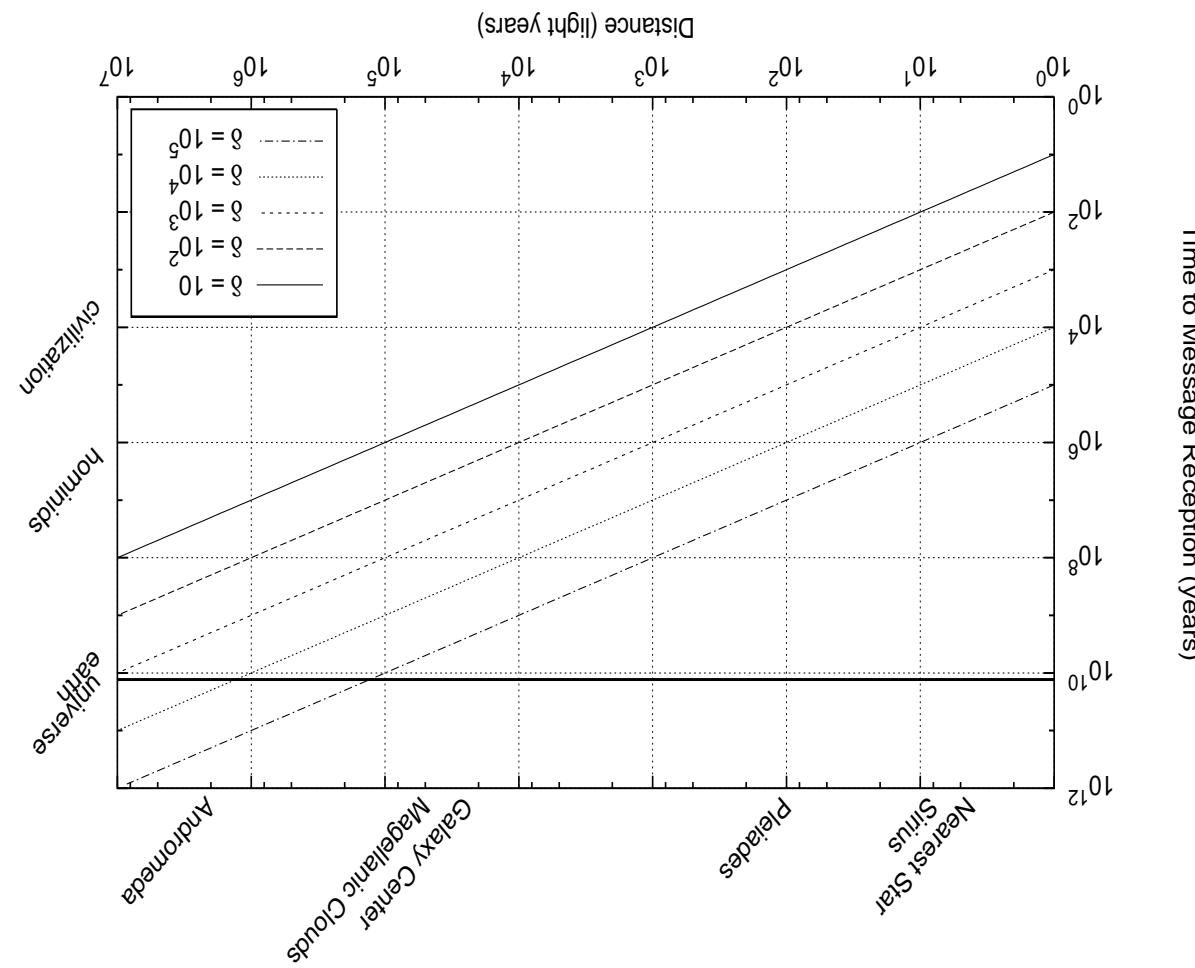
Same but with gain.



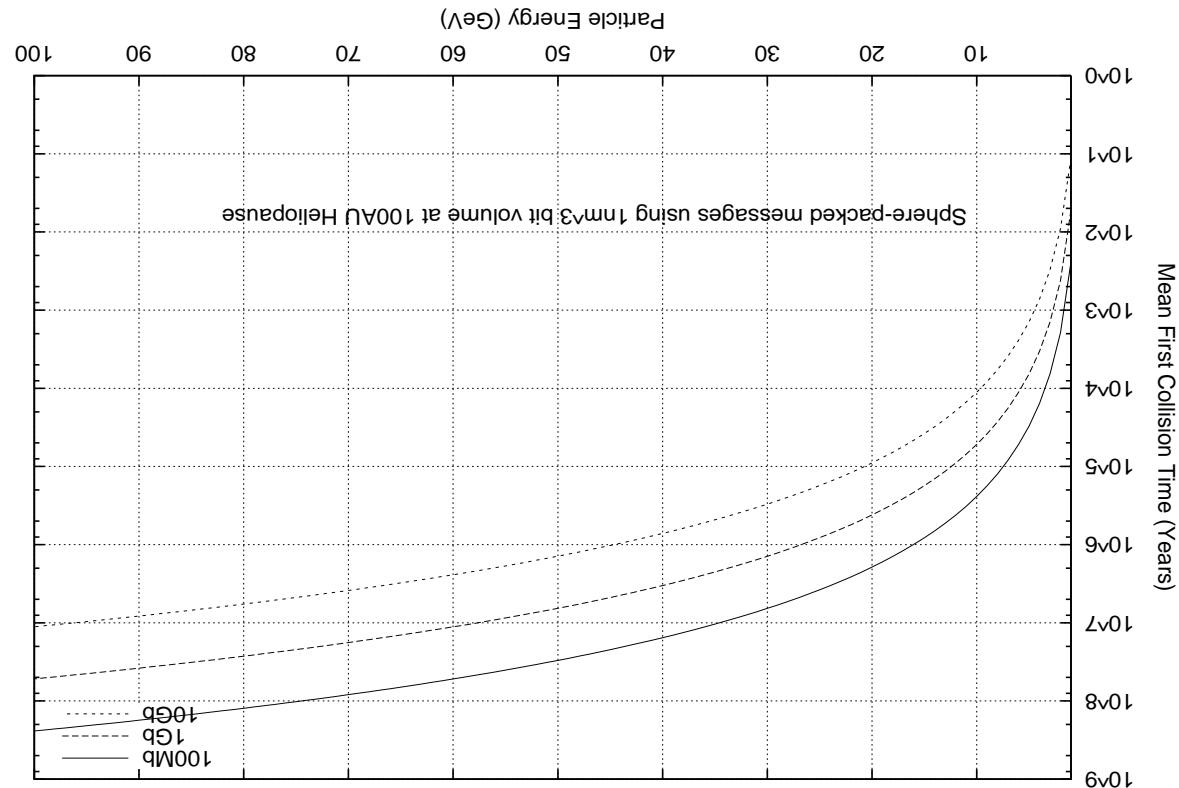
Many Messages with Gain: terrestrial

- CAVEMAT: E^* goes down as δ^{-2}
- E^* at $\delta = 100: 2.5 \times 10^{-12} \text{ J bit}^{-1}$
- Replication: $\approx 10^{-17} \text{ J bit}^{-1}$
- $8 \times 10^{-20} \text{ J per ATP}$
- 60000 ATP/second for 20 minutes: 4850 Kbase of E-coli
- Empirical energy calc:
 - Landauer said it can be arbitrarily fast
- Mass Incription/Readout Time
 - Landauer said it can be reversible
- Mass Incription/Readout Energy
 - Landauer said it can be irreversable

Inscription Energy/Speed Issues



Delivery Times



Message Integrity - High Energy Insights

- Spallation
- Dislocation
- Ion tracks
- Heating (diffusion)
- Large (or small) unseen obstacles

Damage Mechanisms

Hard to Speculate (so won't ... yet)

- Onward toward Lunatic fringe
- Probe (Bracewell)
- Embedded dust (comet)?
- Dust?
- Big rock?

Delivery Methods

The epiphany is REAL even for inefficient delivery!

- 0.5 kg of 20 lb paper at 600 dots/inch
- 5.5×10^{-12} grams of RNA
- D_4 isotropic break-even (1m^2 radius aperture): $\approx 10^{10}$ bits
- 20 miles: 1.2×10^8 Joules
- 20 miles/gallon
- Gallon of gasoline: 1.2×10^8 Joules

Epiphany Revisited, Just for Fun!

- Learn more: <http://www.winlab.rutgers.edu/~crose/cgi-bin/cosmic.html>
 - Interstellar: dust/pebbles/rocks, not spectrum for SETI
 - Terrestrial: maybe some, maybe none
- Practical value
 - Delays reasonable (Boston/NYC: ≈ 270 seconds ballistically)
 - Pebble net throughput scales well
 - Pebbles are non-interfering
- Network Issues
 - Inscribed mass messaging might often be PREFERRED
 - Inscribed mass messaging is NOT ridiculous

PUNCHINES